

$$(i \not{\partial} - m) \psi = 0$$

$$\psi(x) = \int \frac{d^4 k}{(2\pi)^4} \hat{\psi}(k) e^{-i k x}$$

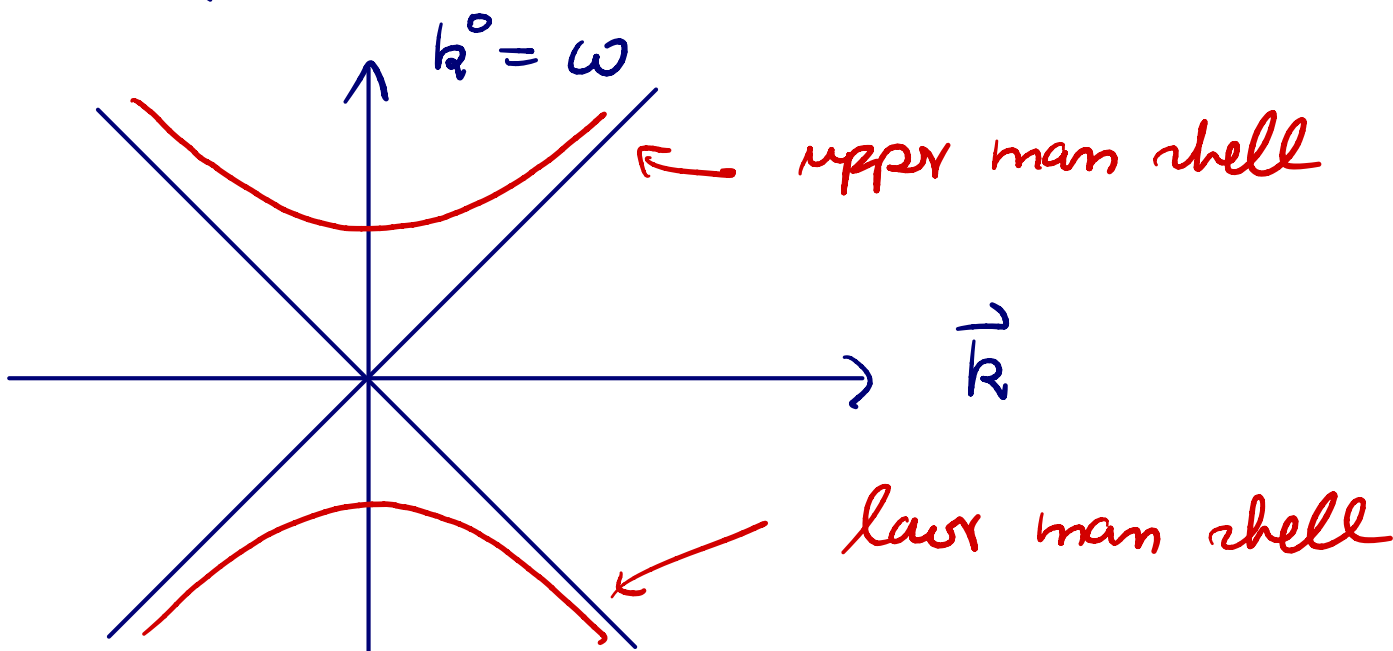
$\langle k, x \rangle$

$$\Leftrightarrow (k - m) \hat{\psi}(k) = 0 \quad | \times (k + m)$$

$$(k + m)(k - m) = (k^2 - m^2) \mathbb{1} \quad (k^2 = \langle k, k \rangle)$$

$$\Rightarrow (k^2 - m^2) \hat{\psi}(k) = 0$$

$$\omega^2 - |\vec{k}|^2 - m^2 = 0$$

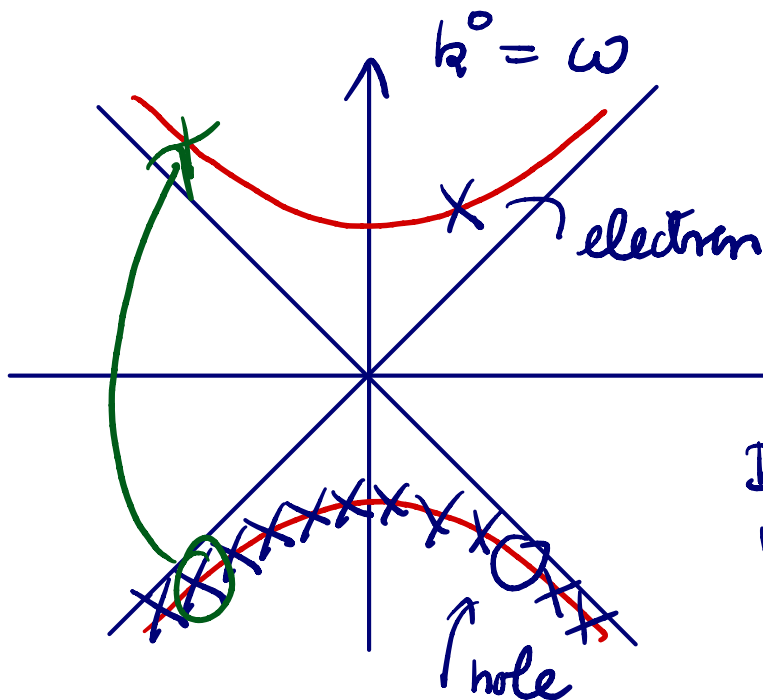


$\hat{\psi}$ is a tempered distribution
 $\psi(x)$ will typically be smooth
 general solution

$$\hat{\psi}(k) = (k + m) \delta(k^2 - m^2) \hat{\phi}(k)$$

$$E = \hbar \omega = \hbar \omega \quad (\text{here } \hbar = 1)$$

Dirac sea



in vacuum, all the states of negative energy are filled

Dirac sea, invisible because it is homogeneous and isotropic

gives rise to anti-matter and pair creation

The hole appears as a particle of positive energy
positron but positive charge

basic drawbacks:

- sea has an infinite charge density
- infinite negative energy density

in modern QFT the Dirac sea no longer appears
because of a suitable "redefinition" of the vacuum.

starting point of causal fermion systems:

Take the Dirac sea picture seriously,
but formulate other equations which do not
suffer from the above drawbacks.