# The Dirac equation and its separation in the Reissner-Nordstroem geometry

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#### Overview

- Reissner-Nordstroem Metric
- Tetrad-Formalism
- Newman-Penrose-Formalism
- The Spin Frame
- Spin Coefficients and Dirac Equation
- Separation of the Dirac Equation

#### Reissner-Nordstroem Metric Basics

- Asymptotically flat, Lorentzian 4-Mfld. (M,g) w/ topology  $S^2 imes \mathbb{R}^2$
- g stationary, radial sym. w/ signature (+, -, -, -)

• 
$$g = \frac{\Delta}{r^2} \mathrm{d}t \otimes \mathrm{d}t - \frac{r^2}{\Delta} \mathrm{d}r \otimes \mathrm{d}r - r^2 \mathrm{d}\theta \otimes \mathrm{d}\theta - r^2 \sin\theta \mathrm{d}\phi \otimes \mathrm{d}\phi$$

• 
$$\Delta(r) = r^2 - 2Mr + Q^2$$
, two roots at  $r_\pm = M \pm \sqrt{M^2 - Q^2}$  when  $Q < M$ 

## Penrose Diagram



# Eddington-Finkelstein-Coord

• Get rid of coord. singularities at  $r_{\pm}$  by using tortoise coord.  $r_{*}$ 

• 
$$\frac{\mathrm{d}r_*}{\mathrm{d}r} = \frac{r^2}{\Delta} \Rightarrow r + A \ln(r - r_-) - B \ln(r - r_+) = r_*$$

- Compute E-F-Coord. from tangent vectors associated to principal null geodesics - (png)
- $t = \pm r_* + C$ , with " " ingoing and " + " outgoing rays
- Define new time coord.  $au:=t+r_*-r.$  Then au is a Cauchy time function

#### **Tetrad-Formalism Basics**

- pointwise orthonormal vector fields (VF)
- One VF is time-like, the other three are space-like
- At each point of the space-time (ST) set a basis of four contra-variant vectors  $e^{\mu}_{(a)}$  with  $a, \mu \in \{0, 1, 2, 3\}$
- η<sub>(a)(b)</sub> := e<sup>μ</sup><sub>(a)</sub>e<sub>(b)μ</sub> constant, sym. matrix for lowering tetrad indices
- Project arbitrary tensor, i.e.  $T^{\mu}_{\nu\alpha}$ , onto the tetrad frame:  $T^{(a)}_{(b)(c)} = e^{(a)}_{\mu} e^{\nu}_{(b)} e^{\alpha}_{(c)} T^{\mu}_{\nu\alpha}$

# **Ricci-Rotation** Coefficients

$$\mathcal{Y}_{(\alpha)(b)(c)} = \left( \nabla_{\mu} e_{(b)\overline{\mu}} \right) e_{(c)}^{\mu} e_{(a)}^{\overline{\mu}}$$

- Intrinsic derivative is defined:  $T_{(a)|(b)} := T_{(a),(b)} - \eta^{(m)(n)}\gamma_{(n)(a)(b)}T_{(m)}$
- $\gamma_{(a)(b)(c)}$  describes the Ricci rotation coefficients, also often called scalar fields
- When implying orthogonal tetrad in local Minkowski space  $\gamma_{(a)(b)(c)} = -\gamma_{(b)(a)(c)}$ . Therefore, 24 independent, real scalar fields
- Evaluation of Ricci rotation coeff. does not involve the evaluation of covariant derivatives.

#### Newman-Penrose-Formalism Basics

• Four null vectors  $\{l^{\mu}, n^{\mu}, m^{\mu}, \bar{m}^{\mu}\}$  with normalization conditions  $l^{\mu}n_{\mu} = 1$  and  $m^{\mu}\bar{m}_{\mu} = -1$ 

• 
$$e^{\mu}_{(0)} = l^{\mu}$$
,  $e^{\mu}_{(1)} = n^{\mu}$ ,  $e^{\mu}_{(2)} = m^{\mu}$  and  $e^{\mu}_{(3)} = ar{m}^{\mu}$ 

Define 12 complex spin coefficients in the NP-Formalism due to the Ricci rotation coefficients.
 κ = γ<sub>(2)(0)(0)</sub> = ½(λ<sub>(2)(0)(0)</sub> + λ<sub>(0)(2)(0)</sub> − λ<sub>(0)(0)(2)</sub>)

• 
$$\lambda_{(a)(b)(c)} := e_{(b)\mu,\nu} (e^{\mu}_{(a)} e^{\nu}_{(c)} - e^{\nu}_{(a)} e^{\mu}_{(c)})$$

## Lorentz Transformations

- Most time ST has not enough local structure to define four vectors for a complete tetrad.
- In NP-Formalism: I<sup>μ</sup> and n<sup>μ</sup> are determined by pngs. m<sup>μ</sup> and m<sup>μ</sup> are unit space-like VF, orthogonal to itself, I<sup>μ</sup> and n<sup>μ</sup>.
- Therefore, it exists a two dim. gauge freedom which is described by the two param. subgroup of the Lorentz Group sometimes denoted by **rotations of class III**, leaving the direction of  $I^{\mu}$  and  $n^{\mu}$  unchanged.
- Generated by boosts  $I^{\mu} \longrightarrow r I^{\mu}$ ,  $n^{\mu} \longrightarrow r^{-1} n^{\mu}$  and rotations  $m^{\mu} \longrightarrow e^{i\alpha} m^{\mu}$

- ST of GR is locally Minkowskian. Therefore, define locally tetrad basis for spinors  $\zeta^A_{(a)}$  and  $\zeta^{A'}_{(a')}$
- A and A' are spinor indices where A corresponds to the fundamental and A' to the anti-fundamental representation
- Sometimes convenient use special symbols  $\zeta^A_{(0)}=\sigma^A$  and  $\zeta^A_{(1)}=\iota^A$
- $\epsilon_{AB}$  skew-symmetric metric:  $\epsilon_{AB}\zeta^{A}_{(a)}\zeta^{B}_{(b)} = \zeta_{(a)B}\zeta^{B}_{(b)} = \epsilon_{(a)(b)}$

## Generalized Pauli Spin-Matrices

• Spinors and their complex conjugate determine  $\{I^{\mu}, n^{\mu}, m^{\mu}, \bar{m}^{\mu}\}$  by the correspondence

• 
$$I^{\mu} \leftrightarrow \sigma^A \bar{\sigma}^{B'}$$
,  $n^{\mu} \leftrightarrow \iota^A \bar{\iota}^{B'} m^{\mu} \leftrightarrow \sigma^A \bar{\iota}^{B'}$  and  $\bar{m}^{\mu} \leftrightarrow \iota^A \bar{\sigma}^{B'}$ 

 Due to this representation one can define the hermitian matrices (generalized Pauli spin-matrices) σ<sup>μ</sup><sub>AB'</sub>:

• 
$$\sigma^{\mu}_{AB'} = \frac{1}{\sqrt{2}} \begin{bmatrix} I^{\mu} & m^{\mu} \\ \bar{m}^{\mu} & n^{\mu} \end{bmatrix}, \qquad \sigma_{AB'\mu} = \frac{1}{\sqrt{2}} \begin{bmatrix} I_{\mu} & -m_{\mu} \\ -\bar{m}_{\mu} & n_{\mu} \end{bmatrix}$$

• This null tetrad fulfils the normalization conditions:

$$I^{\mu}n_{\mu} = \underbrace{\sigma^{\mu}_{AB'}\sigma^{AB'\mu}}_{\delta^{\mu}_{\mu}=1} \sigma^{A}\bar{\sigma}^{B'}\iota_{A}\bar{\iota}_{B'} = 1$$
$$m^{\mu}\bar{m}_{\mu} = \underbrace{\sigma^{\mu}_{AB'}\sigma^{AB'\mu}}_{\delta^{\mu}_{\mu}=1} \sigma^{A}\bar{\iota}^{B'}\iota_{A}\bar{\sigma}_{B'} = -1$$

• Dyad basis determine four null vectors which can be used as a basis for the NP-Formalism

# Dyad Spin Coefficients

 Covariant derivative of a spinor field satisfies the Leibnitz rule, is a real operator and based on correspondences:

$$abla_{\mu} \leftrightarrow 
abla_{AB'}, \qquad 
abla_{\mu} X_{\nu} \leftrightarrow 
abla_{AB'} X_{CD'}$$

 Define analogous an intrinsic derivative for the dyad components ξ<sub>(a)</sub> of a spinor along (a)(b'):

$$\begin{aligned} \xi_{(c)|(a)(b')} &= (\nabla_{AB'}\xi_C\zeta_{(c)}^C)\zeta_{(a)}^A\zeta_{(b)}^{B'} \\ \Leftrightarrow \xi_{(c)|AB'} &= (\nabla_{AB'}\xi_C\zeta_{(c)}^C) \\ \Rightarrow \xi_{(a)|BC'} &= \xi_{(a),BC'} + \Gamma_{(d)(a)BC'}\xi^{(d)} \end{aligned}$$

- 12 independent complex coefficients
- One can show with one Lemma from Friedman (states an alternative rep. of Γ<sub>(a)(b)CD'</sub>) following theorem:
   Theorem:

It is possible to express the dyad spin coefficients  $\Gamma_{(a)(b)CD'}$  in terms of the covariant derivatives of the basis null vectors  $\{l^{\mu}, n^{\mu}, m^{\mu}, \bar{m}^{\mu}\}$  and therefore show that they are in agreement with the coefficients  $\gamma_{(a)(b)(c)}$ .

$$\int_{0000}^{7} = 4 = \chi_{(2)(6)(6)} = \frac{1}{2} \left( \lambda_{(2)(6)(6)} + \lambda_{(6)(2)(6)} - \lambda_{(6)(6)(2)} \right)$$

# Spin Coefficients in Reissner-Nordstroem

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- Use tangent vectors associated to the png to define null tetrad
- Let  $(\psi, U)$  be a local param. of M with  $p \in U \subset M$ ,  $\psi(p) = (t, r, \theta, \phi)$ .  $\{\partial_t, \partial_r, \partial_\theta, \partial_\phi\}$  is the induced canonical basis of  $T_pM$  and  $\{dt, dr, d\theta, d\phi\}$  of  $T_pM^*$ .

$$I^{\mu} = \frac{1}{|\Delta|} (r^{2} \partial_{t} + \Delta \partial_{r}), \quad n^{\mu} = \frac{\operatorname{sign}(\Delta)}{2r^{2}} (r^{2} \partial_{t} - \Delta \partial_{r})$$
$$m^{\mu} = \frac{1}{\sqrt{2}r} (\partial_{\theta} + i \operatorname{csc}(\theta) \partial_{\phi}), \quad \bar{m}^{\mu} = \frac{1}{\sqrt{2}r} (\partial_{\theta} - i \operatorname{csc}(\theta) \partial_{\phi})$$

#### Spin Coefficients in Reissner-Nordstroem II

- Class III Lorentz transformation:  $r=\sqrt{rac{|\Delta|}{2r^2}}$  and lpha=0
- Re-write in Eddington-Finkelstein-Coordinates  $au = t + r_* r$
- Additionally Class III Lorentz transformation:  $r^{'} = \frac{\sqrt{|\Delta|}}{r_{+}}$  and  $\alpha^{'} = 0$

$$I'' = \frac{1}{\sqrt{2}rr_{+}} \left[ \left( 2r^{2} - \Delta \right) \partial_{\tau} + \Delta \partial_{r} \right]$$
$$I''_{D} = \frac{1}{\sqrt{2}rr_{+}} \left[ \Delta d\tau + \left( \Delta - 2r^{2} \right) dr \right]$$

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# Spin Coefficients in Reissner-Nordstroem III

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 Compute Spin coefficients in NP-Formalism for Reissner-Nordstroem ST. Six coefficients are distinct from zero.

$$\begin{aligned} \pi'' &= \tau'' = \kappa'' = \sigma'' = \nu'' = \lambda'' = 0\\ \alpha'' &= -\beta'' = \frac{1}{2^{3/2}} \frac{\cot(\theta)}{r}, \qquad \gamma'' = -\frac{r_+}{2^{3/2}} \frac{1}{r^2}\\ \epsilon'' &= \frac{1}{2^{3/2}r_+} \left(1 - \frac{Q^2}{r^2}\right), \qquad \rho'' = -\frac{1}{\sqrt{2}r_+} \frac{\Delta}{r^2}\\ \mu'' &= -\frac{r_+}{\sqrt{2}} \frac{1}{r^2} \end{aligned}$$

• Consistent with results from C. Röken for a Kerr ST in the limit  $a \longrightarrow 0$ .

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# General Dirac Equation

- Got spin coefficients in the NP-Formalism. But where is the connection with the Dirac-Eq.?
- (M,g) is an arbitrary 4-dim. curved ST,  $S = P \times_{\tau} \Delta_4$  the associated spinor bundle,  $\tau : \text{Spin}(4) \longrightarrow \text{GL}(\Delta_4)$  and (P, F) the spin structure.  $\Psi \in \Gamma(S) \simeq \mathbb{C}^4$  is a Dirac four spinor and m its invariant fermion rest mass.

• 
$$\left[\gamma^{\mu}\nabla^{\mathcal{S}}_{\mu} + im\right]\Psi(x^{\mu}) = 0 \text{ with } \{\gamma^{\mu}, \gamma^{\nu}\} = 2g^{\mu\nu}\mathbb{1}^{\mathbb{C}}_{4\times 4}$$

•  $\gamma^{\mu}$  general relativist Dirac matrices and  $\nabla^{S}$  being the metric connection in the spinor bundle

# General Dirac Equation in Spinor Rep.

• Two spinor rep. of 
$$\Psi = \begin{pmatrix} P^A \\ \bar{Q}_{B'} \end{pmatrix}$$
  
•  $\gamma^{\mu} = \gamma^{\mu} = \sqrt{2} \begin{bmatrix} 0^{\mathbb{C}}_{2\times 2} & \sigma^{\mu AB'} \\ \sigma^{\mu}_{AB'} & 0^{\mathbb{C}}_{2\times 2} \end{bmatrix}$  with  $\sigma^{\mu}_{AB'}$  the generalized Pauli matrices

$$\nabla_{AB'}P^{A} + \frac{im}{\sqrt{2}}\bar{Q}^{C'}\epsilon_{C'B'} = 0$$
$$\nabla_{AB'}Q^{A} + \frac{im}{\sqrt{2}}\bar{P}^{C'}\epsilon_{C'B'} = 0$$

# General Dirac Equation in Spinor Rep. II

• Example calculation for 
$$B' = 0$$
:  $\nabla_{A0'}P^A + \frac{im}{\sqrt{2}}\bar{Q}_{a}^{1'}\epsilon_{1'0'} = 0$   
 $\sum_{a=-1}^{\infty} \sqrt{2}\bar{Q}_{a}^{1'}\epsilon_{1'0'} = 0$   
 $\sum_{a=-1}^{\infty} \sqrt{2}\bar{Q}_{a}^{1'}\epsilon_{1'0'} = 0$   
 $\sum_{a=-1}^{\infty} \sqrt{2}\bar{Q}_{a}^{1'}\epsilon_{1'0'} = 0$   
 $\sum_{a=-1}^{\infty} \sqrt{2}\bar{Q}_{a}^{1'}\epsilon_{1'$ 

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General Dirac Equation in Spinor Rep. II

$$(D+E-S)P^{0} + (\overline{\delta}+\overline{1}\overline{1}-a)P^{1} = \frac{im}{\sqrt{2}}\overline{G}^{1}$$

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# General Dirac Equation in Spinor Rep. III

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• Substituting: 
$$\mathcal{F}_0=P^0$$
,  $\mathcal{F}_1=P^1$ ,  $\mathcal{G}_0=ar{Q}^{1'}$  and  $\mathcal{G}_1=-ar{Q}^{0'}$ 

$$(D + \epsilon - \rho)\mathcal{F}_{0} + (\bar{\delta} + \pi - \alpha)\mathcal{F}_{1} = \frac{im}{\sqrt{2}}\mathcal{G}_{0}$$
$$(\delta + \beta - \tau)\mathcal{F}_{0} + (\Delta + \mu - \gamma)\mathcal{F}_{1} = \frac{im}{\sqrt{2}}\mathcal{G}_{1}$$
$$(D + \bar{\epsilon} - \bar{\rho})\mathcal{G}_{1} - (\delta + \bar{\pi} - \bar{\alpha})\mathcal{G}_{0} = \frac{im}{\sqrt{2}}\mathcal{F}_{1}$$
$$(\Delta + \bar{\mu} - \bar{\gamma})\mathcal{G}_{0} - (\bar{\delta} + \bar{\beta} - \bar{\tau})\mathcal{G}_{1} = \frac{im}{\sqrt{2}}\mathcal{F}_{0}$$

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#### Separation of the Dirac Equation

• Ansatz: 
$$\mathcal{F}_i = e^{i(\omega \tau + k\phi)} \mathcal{H}_i(r, \theta)$$
 and  $\mathcal{G}_i = e^{i(\omega \tau + k\phi)} \mathcal{J}_i(r, \theta)$ 

Define:

$$\chi(r, Q) := \frac{1}{2r^2} (Q^2 + r(3r - 4M))$$
$$\mathcal{L}_n := \partial_\theta + n\cot(\theta) + k\csc(\theta)$$

• Looking at the first ODE:  $\frac{1}{r_{+}} \left[ i\omega(2r^{2} - \Delta) + \Delta\partial_{r} + \chi \right] \mathcal{H}_{0}(r, \theta) + \mathcal{L}_{1/2}\mathcal{H}_{1} = imr\mathcal{J}_{0}(r, \theta)$ 

## Separation of the Dirac Equation II

• Separation Ansatz:

$$\begin{aligned} \mathcal{H}_0 &= R_+(r)S_+(\theta), \qquad \mathcal{H}_1 &= R_-(r)S_-(\theta) \\ \mathcal{J}_0 &= R_-(r)S_+(\theta), \qquad \mathcal{J}_1 &= R_+(r)S_-(\theta) \end{aligned}$$

• These ansatz results in following equation:  

$$\left(\frac{1}{r_{+}} \left[i\omega(2r^{2}-\Delta)+\Delta\partial_{r}+\chi(r,\theta)\right]R_{+}(r)-imrR_{-}(r)\right)S_{+}(\theta) + \mathcal{L}_{\frac{1}{2}}S_{-}(\theta)R_{-}(r) = 0$$

• Therefore, following must be satisfied:

$$-\lambda S_{+}(\theta) = \mathcal{L}_{\frac{1}{2}}S_{-}(\theta)$$
$$\lambda R_{-}(r) = \frac{1}{r_{+}} \left[i\omega(2r^{2} - \Delta) + \Delta\partial_{r} + \chi(r,\theta)\right]R_{+}(r) - imrR_{-}(r)$$

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with  $\lambda$  being the constant of separation.

#### Separation of the Dirac Equation III

- Analogue for the other three equations. Resulting in four ODEs:
- Two radial equations:

$$\begin{bmatrix} \Delta \partial_r + \chi(r, Q) + i\omega(2r^2 - \Delta) \end{bmatrix} R_+(r) = r_+(\lambda + imr)R_-(r)$$
$$r_+(\partial_r + \frac{1}{r} - i\omega)R_-(r) = (\lambda - imr)R_+(r)$$

• Two angular equations:

$$\mathcal{L}_{rac{1}{2}}S_{-}( heta) = -\lambda S_{+}( heta)$$
 $\mathcal{L}_{rac{1}{2}}^{\dagger}S_{+}( heta) = \lambda S_{-}( heta)$ 

- R. G. S. K. Chakrabarti and C. bin Liang, "Timelike curves of limited acceleration in general relativity," *Journal of Mathematical Physics*, vol. 24, no. 3, pp. 597–598, 1983.
- S. Chandrasekhar, *The Mathematical Theory of Black Holes*. Walton Street, Oxford OX2 6DP: Oxford University Press, 1 ed., 1983.
- [3] R. P. R. Geroch and A. Held, "A space-time calculus based on paris of null directions," *Journal of Mathematical Physics*, vol. 14, no. 7, pp. 874–881, 1973.
- [4] Z. Ahsan, The Potential of Fields in Einstein's Theory of Gravitation.
  152 Beach Road, 21-01/04 Gateway East, Singapore 189721, Singapore: Springer, 1 ed., 2019.

- [5] C. Röken, "The massive dirac equation in kerr geometry: separability in eddington-finkelstein-type coordinates and asymptotics," *General Relativity and Gravitation*, vol. 49, Feb 2017.
- [6] "Spinors in 1+3 dimensions," 2015.

"Online erhältlich unter https:

//www.physik.uni-muenchen.de/lehre/vorlesungen/
wise\_15\_16/tv\_qm2/vorlesung/Spinors-4D.pdf; abgerufen
am 04. Januar 2021.".